

## VU Research Portal

### The Q2 dependence of the generalized Gerasimov-Drell-Hearn integral for the proton.

Airapetian, A.; Amarian, M.; Aschenauer, E.C.; Bulten, H.J.; Ferro-Luzzi, M.M.E.; Garutti, E.; Heesbeen, D.; Hesselink, W.H.A.; Hoffman-Rothe, P.; Laziev, A.; van den Brand, J.F.J.; van der Steenhoven, G.; van Hunen, J.J.; Visser, J.

**published in**

Physics Letters B  
2000

**DOI (link to publisher)**

[10.1016/S0370-2693\(00\)01111-4](https://doi.org/10.1016/S0370-2693(00)01111-4)

**document version**

Publisher's PDF, also known as Version of record

[Link to publication in VU Research Portal](#)

**citation for published version (APA)**

Airapetian, A., Amarian, M., Aschenauer, E. C., Bulten, H. J., Ferro-Luzzi, M. M. E., Garutti, E., Heesbeen, D., Hesselink, W. H. A., Hoffman-Rothe, P., Laziev, A., van den Brand, J. F. J., van der Steenhoven, G., van Hunen, J. J., & Visser, J. (2000). The Q2 dependence of the generalized Gerasimov-Drell-Hearn integral for the proton. *Physics Letters B*, 494, 1-8. [https://doi.org/10.1016/S0370-2693\(00\)01111-4](https://doi.org/10.1016/S0370-2693(00)01111-4)

**General rights**

Copyright and moral rights for the publications made accessible in the public portal are retained by the authors and/or other copyright owners and it is a condition of accessing publications that users recognise and abide by the legal requirements associated with these rights.

- Users may download and print one copy of any publication from the public portal for the purpose of private study or research.
- You may not further distribute the material or use it for any profit-making activity or commercial gain
- You may freely distribute the URL identifying the publication in the public portal ?

**Take down policy**

If you believe that this document breaches copyright please contact us providing details, and we will remove access to the work immediately and investigate your claim.

**E-mail address:**

[vuresearchportal.ub@vu.nl](mailto:vuresearchportal.ub@vu.nl)

# The $Q^2$ -dependence of the generalised Gerasimov–Drell–Hearn integral for the proton

HERMES Collaboration

A. Airapetian<sup>ae</sup>, N. Akopov<sup>ae</sup>, I. Akushevich<sup>g</sup>, M. Amarian<sup>w,z,ae</sup>, J. Arrington<sup>b</sup>,  
 E.C. Aschenauer<sup>g,m,w</sup>, H. Avakian<sup>k</sup>, R. Avakian<sup>ae</sup>, A. Avetissian<sup>ae</sup>, E. Avetissian<sup>ae</sup>,  
 P. Bailey<sup>o</sup>, B. Bains<sup>o</sup>, C. Baumgarten<sup>u</sup>, M. Beckmann<sup>l</sup>, S. Belostotski<sup>x</sup>,  
 S. Bernreuther<sup>i</sup>, N. Bianchi<sup>k</sup>, H. Böttcher<sup>g</sup>, A. Borissov<sup>f,n</sup>, M. Bouwhuis<sup>o</sup>, J. Brack<sup>e</sup>,  
 S. Brauksiepe<sup>l</sup>, B. Braun<sup>i,u</sup>, W. Brückner<sup>n</sup>, A. Brüll<sup>n,r</sup>, P. Budz<sup>i</sup>, H.J. Bulten<sup>q,w,ad</sup>,  
 G.P. Capitani<sup>k</sup>, P. Carter<sup>d</sup>, P. Chumney<sup>v</sup>, E. Cisbani<sup>z</sup>, G.R. Court<sup>p</sup>, P.F. Dalpiaz<sup>j</sup>,  
 R. De Leo<sup>c</sup>, L. De Nardo<sup>a</sup>, E. De Sanctis<sup>k</sup>, D. De Schepper<sup>b,r</sup>, E. Devitsin<sup>t</sup>,  
 P.K.A. de Witt Huberts<sup>w</sup>, P. Di Nezza<sup>k</sup>, V. Djordjadze<sup>g</sup>, M. Düren<sup>i</sup>, A. Dvoredsky<sup>d</sup>,  
 G. Elbakian<sup>ae</sup>, J. Ely<sup>e</sup>, A. Fantoni<sup>k</sup>, A. Fechtchenko<sup>h</sup>, M. Ferro-Luzzi<sup>w</sup>, K. Fiedler<sup>i</sup>,  
 B.W. Filippone<sup>d</sup>, H. Fischer<sup>l</sup>, B. Fox<sup>e</sup>, J. Franz<sup>l</sup>, S. Frullani<sup>z</sup>, Y. Gärber<sup>g</sup>,  
 F. Garibaldi<sup>z</sup>, E. Garutti<sup>w</sup>, G. Gavrilo<sup>x</sup>, V. Gharibyan<sup>ae</sup>, A. Golendukhin<sup>f,u,ae</sup>,  
 G. Graw<sup>u</sup>, O. Grebeniouk<sup>x</sup>, P.W. Green<sup>a,ab</sup>, L.G. Greeniaus<sup>a,ab</sup>, A. Gute<sup>i</sup>, W. Haeberli<sup>q</sup>,  
 M. Hartig<sup>ab</sup>, D. Hasch<sup>g,k</sup>, D. Heesbeen<sup>w</sup>, F.H. Heinsius<sup>l</sup>, M. Henoch<sup>i</sup>,  
 R. Hertenberger<sup>u</sup>, W. Hesselink<sup>w</sup>, P. Hoffmann-Rothe<sup>w</sup>, G. Hofman<sup>e</sup>, Y. Holler<sup>f</sup>,  
 R.J. Holt<sup>o</sup>, B. Hommez<sup>m</sup>, G. Iarygin<sup>h</sup>, M. Iodice<sup>z</sup>, A. Izotov<sup>x</sup>, H.E. Jackson<sup>b</sup>,  
 A. Jgoun<sup>x</sup>, P. Jung<sup>g</sup>, R. Kaiser<sup>g,aa,ab</sup>, J. Kanesaka<sup>ac</sup>, E. Kinney<sup>e</sup>, A. Kisselev<sup>x</sup>,  
 P. Kitching<sup>a</sup>, H. Kobayashi<sup>ac</sup>, N. Koch<sup>i</sup>, K. Königsmann<sup>l</sup>, H. Kolster<sup>u</sup>, V. Korotkov<sup>g</sup>,  
 E. Kotik<sup>a</sup>, V. Kozlov<sup>t</sup>, V.G. Krivokhijine<sup>h</sup>, G. Kyle<sup>v</sup>, L. Lagamba<sup>c</sup>, A. Laziev<sup>w</sup>,  
 P. Lenisa<sup>j</sup>, T. Lindemann<sup>f</sup>, W. Lorenzon<sup>s</sup>, N.C.R. Makins<sup>b,o</sup>, J.W. Martin<sup>r</sup>,  
 H. Marukyan<sup>ae</sup>, F. Masoli<sup>j</sup>, M. McAndrew<sup>p</sup>, K. McIlhany<sup>d,r</sup>, R.D. McKeown<sup>d</sup>,  
 F. Menden<sup>l,ab</sup>, A. Metz<sup>u</sup>, N. Meyners<sup>f</sup>, O. Mikloukho<sup>x</sup>, C.A. Miller<sup>a,ab</sup>, R. Milner<sup>r</sup>,  
 V. Mitsyn<sup>h</sup>, V. Muccifora<sup>k</sup>, R. Mussa<sup>j</sup>, A. Nagaitsev<sup>h</sup>, E. Nappi<sup>c</sup>, Y. Naryshkin<sup>x</sup>,  
 A. Nass<sup>i</sup>, K. Negodaeva<sup>g</sup>, W.-D. Nowak<sup>g</sup>, T.G. O'Neill<sup>b</sup>, R. Openshaw<sup>ab</sup>, J. Ouyang<sup>ab</sup>,  
 B.R. Owen<sup>o</sup>, S.F. Pate<sup>r,v</sup>, S. Potashov<sup>t</sup>, D.H. Potterveld<sup>b</sup>, G. Rakness<sup>e</sup>, V. Rappoport<sup>x</sup>,  
 R. Redwine<sup>r</sup>, D. Reggiani<sup>j</sup>, A.R. Reolon<sup>k</sup>, R. Ristinen<sup>e</sup>, K. Rith<sup>i</sup>, D. Robinson<sup>o</sup>,  
 M. Ruh<sup>l</sup>, D. Ryckbosch<sup>m</sup>, Y. Sakemi<sup>ac</sup>, I. Savin<sup>h</sup>, C. Scarlett<sup>s</sup>, A. Schäfer<sup>y</sup>, C. Schill<sup>l</sup>,  
 F. Schmidt<sup>i</sup>, G. Schnell<sup>v</sup>, K.P. Schüller<sup>f</sup>, A. Schwind<sup>g</sup>, J. Seibert<sup>l</sup>, B. Seitz<sup>a</sup>,  
 T.-A. Shibata<sup>ac</sup>, T. Shin<sup>r</sup>, V. Shutov<sup>h</sup>, C. Simani<sup>j</sup>, A. Simon<sup>l,v</sup>, K. Sinram<sup>f</sup>,  
 E. Steffens<sup>h</sup>, J.J.M. Steijger<sup>w</sup>, J. Stewart<sup>p,ab</sup>, U. Stösslein<sup>g</sup>, K. Suetsugu<sup>ac</sup>, M. Sutter<sup>r</sup>,

H. Tallini<sup>p</sup>, S. Taroian<sup>ae</sup>, A. Terkulov<sup>t</sup>, S. Tessarin<sup>j</sup>, E. Thomas<sup>k</sup>, B. Tipton<sup>r,d</sup>,  
 M. Tytgat<sup>m</sup>, G.M. Urciuoli<sup>z</sup>, J.F.J. van den Brand<sup>w,ad</sup>, G. van der Steenhoven<sup>w</sup>,  
 R. van de Vyver<sup>m</sup>, J.J. van Hunen<sup>w</sup>, M.C. Vetterli<sup>aa,ab</sup>, V. Vikhrov<sup>x</sup>, M.G. Vincter<sup>ab,a,\*</sup>,  
 J. Visser<sup>w</sup>, E. Volk<sup>n</sup>, C. Weiskopf<sup>i</sup>, J. Wendland<sup>aa,ab</sup>, J. Wilbert<sup>i</sup>, T. Wise<sup>q</sup>, S. Yen<sup>ab</sup>,  
 S. Yoneyama<sup>ac</sup>, H. Zohrabian<sup>ae</sup>

<sup>a</sup> Department of Physics, University of Alberta, Edmonton, AL T6G 2J1, Canada

<sup>b</sup> Physics Division, Argonne National Laboratory, Argonne, IL 60439-4843, USA

<sup>c</sup> Istituto Nazionale di Fisica Nucleare, Sezione di Bari, 70124 Bari, Italy

<sup>d</sup> W.K. Kellogg Radiation Laboratory, California Institute of Technology, Pasadena, CA 91125, USA

<sup>e</sup> Nuclear Physics Laboratory, University of Colorado, Boulder, CO 80309-0446, USA

<sup>f</sup> DESY, Deutsches Elektronen Synchrotron, 22603 Hamburg, Germany

<sup>g</sup> DESY Zeuthen, 15738 Zeuthen, Germany

<sup>h</sup> Joint Institute for Nuclear Research, 141980 Dubna, Russia

<sup>i</sup> Physikalisches Institut, Universität Erlangen-Nürnberg, 91058 Erlangen, Germany

<sup>j</sup> Istituto Nazionale di Fisica Nucleare, Sezione di Ferrara and Dipartimento di Fisica, Università di Ferrara, 44100 Ferrara, Italy

<sup>k</sup> Istituto Nazionale di Fisica Nucleare, Laboratori Nazionali di Frascati, 00044 Frascati, Italy

<sup>l</sup> Fakultät für Physik, Universität Freiburg, 79104 Freiburg, Germany

<sup>m</sup> Department of Subatomic and Radiation Physics, University of Gent, 9000 Gent, Belgium

<sup>n</sup> Max-Planck-Institut für Kernphysik, 69029 Heidelberg, Germany

<sup>o</sup> Department of Physics, University of Illinois, Urbana, IL 61801, USA

<sup>p</sup> Physics Department, University of Liverpool, Liverpool L69 7ZE, United Kingdom

<sup>q</sup> Department of Physics, University of Wisconsin-Madison, Madison, WI 53706, USA

<sup>r</sup> Laboratory for Nuclear Science, Massachusetts Institute of Technology, Cambridge, MA 02139, USA

<sup>s</sup> Randall Laboratory of Physics, University of Michigan, Ann Arbor, MI 48109-1120, USA

<sup>t</sup> Lebedev Physical Institute, 117924 Moscow, Russia

<sup>u</sup> Sektion Physik, Universität München, 85748 Garching, Germany

<sup>v</sup> Department of Physics, New Mexico State University, Las Cruces, NM 88003, USA

<sup>w</sup> Nationaal Instituut voor Kernfysica en Hoge-Energiefysica (NIKHEF), 1009 DB Amsterdam, The Netherlands

<sup>x</sup> Petersburg Nuclear Physics Institute, St. Petersburg, Gatchina, 188350 Russia

<sup>y</sup> Institut für Theoretische Physik, Universität Regensburg, 93040 Regensburg, Germany

<sup>z</sup> Istituto Nazionale di Fisica Nucleare, Sezione Sanità and Physics Laboratory, Istituto Superiore di Sanità, 00161 Roma, Italy

<sup>aa</sup> Department of Physics, Simon Fraser University, Burnaby, BC V5A 1S6, Canada

<sup>ab</sup> TRIUMF, Vancouver, BC V6T 2A3, Canada

<sup>ac</sup> Department of Physics, Tokyo Institute of Technology, Tokyo 152, Japan

<sup>ad</sup> Department of Physics and Astronomy, Vrije Universiteit, 1081 HV Amsterdam, The Netherlands

<sup>ae</sup> Yerevan Physics Institute, 375036, Yerevan, Armenia

Received 28 August 2000; accepted 19 September 2000

Editor: W.-D. Schlatter

## Abstract

The dependence on  $Q^2$  (the negative square of the 4-momentum of the exchanged virtual photon) of the generalised Gerasimov–Drell–Hearn integral for the proton has been measured in the range  $1.2 \text{ GeV}^2 < Q^2 < 12 \text{ GeV}^2$  by scattering longitudinally polarised positrons on a longitudinally polarised hydrogen gas target. The contributions of the nucleon-resonance and deep inelastic regions to this integral have been evaluated separately. The latter has been found to dominate for  $Q^2 > 3 \text{ GeV}^2$ , while both contributions are important at low  $Q^2$ . The total integral shows no significant deviation from a  $1/Q^2$  behaviour in the measured  $Q^2$  range, and thus no sign of large effects due to either nucleon-resonance excitations or nonleading twist. © 2000 Elsevier Science B.V. All rights reserved.

\* Corresponding author.

E-mail address: vincter@hermes.desy.de (M.G. Vincter).

PACS: 13.60.Hb; 13.88.+e; 25.20.Dc; 25.30.Fj

Keywords: Deep inelastic scattering; Sum rules; Asymmetries; Photoabsorption

The Gerasimov–Drell–Hearn (GDH) sum rule [1] relates the anomalous contribution  $\kappa$  in the nucleon magnetic moment to an energy-weighted integral of the difference of the nucleon total spin-dependent photoabsorption cross sections:

$$\int_{\nu_0}^{\infty} [\sigma_{1/2}(\nu) - \sigma_{3/2}(\nu)] \frac{d\nu}{\nu} = -\frac{2\pi^2\alpha}{M^2} \kappa^2, \quad (1)$$

where  $\sigma_{1/2(3/2)}$  is the photoabsorption cross section for total helicity of the photon–nucleon system equal to  $1/2$  ( $3/2$ ),  $\nu$  is the photon energy in the target rest frame,  $\nu_0$  is the pion production threshold and  $M$  is the nucleon mass. For the proton ( $\kappa_p = +1.79$ ) the GDH sum rule prediction is  $-204 \mu\text{b}$ . The importance of this sum rule is due to the fact that it is based mostly on very general principles of causality, unitarity, crossing symmetry, and Lorentz and gauge invariance. It has never been directly tested, due to the need for a circularly polarised beam with a longitudinally polarised target, and a wide range of photon energies that has to be covered. There are available several predictions for the contribution of nucleon-resonance excitation to the GDH integral [2], derived from multipole analyses of data for unpolarised single-pion photoproduction, and a prediction for the contribution of high-energy multihadron production [3], based on a multiple-reggeon exchange analysis of deep inelastic asymmetries. The estimate from multipole analysis was confirmed by preliminary results from the GDH experiment at Mainz [4], which covered the photon-energy range from 200 MeV up to 800 MeV.

The GDH integral can be generalised to the case of absorption of polarised transverse virtual photons with squared four-momentum  $-Q^2$  [5]:

$$I_{\text{GDH}}(Q^2) \equiv \int_{\nu_0}^{\infty} [\sigma_{1/2}(\nu, Q^2) - \sigma_{3/2}(\nu, Q^2)] \frac{d\nu}{\nu} \quad (2)$$

$$= 16\pi^2\alpha \int_0^{x_0} \frac{g_1(x, Q^2) - \gamma^2 g_2(x, Q^2)}{Q^2 \sqrt{1 + \gamma^2}} dx \quad (3)$$

$$= \frac{8\pi^2\alpha}{M} \int_0^{x_0} \frac{A_1(x, Q^2) F_1(x, Q^2)}{K} \frac{dx}{x}, \quad (4)$$

where  $g_1$  and  $g_2$  are the polarised structure functions of the nucleon,  $\gamma^2 = Q^2/\nu^2 = (2Mx)^2/Q^2$ ,  $x = Q^2/2M\nu$  and  $x_0 = Q^2/2M\nu_0$ . The quantity  $A_1$  is the longitudinal asymmetry for virtual photoabsorption, while  $F_1$  is the unpolarised structure function of the nucleon. The Gilman notation [6] for the virtual photon flux factor  $K = \nu\sqrt{1 + \gamma^2}$  has been used. It should be noted that elastic scattering occurring at  $x = 1$  does not contribute to the generalised integral. Other generalisations of the GDH integral also have been considered [5]. They differ from the definition given in Eq. (2) by terms in the integral that are proportional to  $\gamma^2$  and which therefore vanish in both the real-photon ( $Q^2 = 0$ ) and the deep inelastic ( $Q^2 \gg 1 \text{ GeV}^2$  and  $\gamma^2 \rightarrow 0$ ) limits. Since  $\gamma^2$  is not small in the nucleon-resonance region and at moderate  $Q^2$  (e.g.  $\gamma^2$  is larger than unity for the  $P_{33}(1232)$ -resonance for  $0.2 \text{ GeV}^2 < Q^2 < 2 \text{ GeV}^2$ ), these generalisations are equivalent for finite values of  $Q^2$  only if the contributions of the nucleon-resonance excitations are small.

The generalisation of the GDH integral to nonzero photon virtuality  $Q^2$  provides a way to study the transition from polarised lepton scattering from the nucleon, which is dominated by deep inelastic scattering (DIS) at large photon–nucleon centre-of-mass energy  $W = \sqrt{M^2 + 2M\nu - Q^2}$ , to the polarised real photon absorption on the nucleon, which is dominated by nucleon-resonance excitation at low  $W$ . In leading twist (e.g. for  $Q^2 \rightarrow \infty$ ), Eq. (3) simplifies: the elastic contribution excluded from the integral is of higher twist, the factor  $1/\sqrt{1 + \gamma^2}$  is  $1 + \mathcal{O}(1/Q^2)$ , and in leading-twist approximation it is generally believed (though not formally proven) that the Burkhardt–Cottingham sum rule  $\int_0^1 g_2(x, Q^2) dx = 0$  holds. This leads to:

$$I_{\text{GDH}}(Q^2)_{\gamma^2 \rightarrow 0} = \frac{16\pi^2\alpha}{Q^2} \Gamma_1, \quad (5)$$

where

$$\Gamma_1 \equiv \int_0^1 g_1(x) dx$$

is predicted to have only the weak  $Q^2$ -dependence due to QCD evolution. As  $Q^2 \rightarrow 0$ ,  $I_{\text{GDH}}(Q^2)$  must change sign in order to reach the negative value predicted by the GDH sum rule at  $Q^2 = 0$ .

Several phenomenological models have been proposed to describe the dependence of the generalised GDH integral on  $Q^2$  [7–11]. Some of these models predict large effects from nucleon-resonance excitation [9,10] or from higher twist [7,11], even for  $Q^2$  up to a few  $\text{GeV}^2$ . Other models based on chiral perturbation theory have been proposed but their application is limited to  $Q^2 \ll 1 \text{ GeV}^2$  [12].

The contribution of the region  $W^2 \geq 3.24 \text{ GeV}^2$  to the GDH integral defined in Eq. (2) was recently measured [13] for the proton and the neutron in the range  $0.8 \text{ GeV}^2 \leq Q^2 \leq 12 \text{ GeV}^2$ , showing that higher-twist effects do not appear to be significant in the measured region. This paper presents a measurement of the contribution of the resonance region to the GDH integral for the proton in a similar  $Q^2$ -range ( $1.2 \text{ GeV}^2 \leq Q^2 \leq 12 \text{ GeV}^2$ ). In combination with the analysis at higher  $W^2$ , this provides the first experimental determination of essentially the complete GDH integral for the proton over a range of  $Q^2$  values.

The measurement was performed in 1997 with a 27.56 GeV beam of longitudinally polarised positrons incident on a longitudinally polarised  $^1\text{H}$  gas target internal to the HERA storage ring at DESY. The beam polarisation was measured continuously using Compton backscattering of circularly polarised laser light [14]. The average beam polarisation for the analysed data was 0.55.

The HERMES polarised target [15] consists of polarised atomic  $^1\text{H}$  gas confined in a storage cell, which is a 40 cm long open-ended thin-walled elliptical tube located on the beam axis inside the HERA vacuum pipe. It is fed by an atomic-beam source of nuclear-polarised hydrogen based on Stern–Gerlach separation [16]. It provides an areal target density of about  $7 \times 10^{13} \text{ atoms/cm}^2$ . The nuclear polarisation of atoms and the atomic fraction are continuously measured with a Breit–Rabi polarimeter and a target gas analyser [17]. The average value of the tar-

get polarisation for the analysed data was 0.88 [18]. The fractional systematic uncertainties of the beam and target polarisations were 3.4% and 4.5%, respectively. The integrated luminosity for this data set was  $70 \text{ pb}^{-1}$ .

Scattered positrons were detected by the HERMES forward spectrometer, which is described in detail elsewhere [19]. The kinematic requirements on the scattered positrons for the analysis in the nucleon-resonance region were:  $1 \text{ GeV}^2 < W^2 < 4.2 \text{ GeV}^2$ ,  $1.2 \text{ GeV}^2 < Q^2 < 12 \text{ GeV}^2$ . After applying data quality criteria, about 0.13 million events were selected.

For all detected positrons the angular resolution was better than 0.6 mrad, the momentum resolution was better than 1.6% aside from bremsstrahlung tails and the  $Q^2$ -resolution was better than 2.2%. The limited  $W^2$ -resolution (about  $840 \text{ MeV}^2$ , or  $\Delta W \simeq 240 \text{ MeV}$ ) in the resonance region did not allow the contributions of the individual nucleon resonances to be distinguished. To evaluate the smearing corrections and the contaminations intruding into the resonance region from the elastic and deep inelastic regions, events were simulated using a Monte-Carlo code that includes elastic, deep inelastic and resonance contributions. The description of the resonance contribution was based on the model of Ref. [20]. The deep inelastic region was modelled using the parameterisation of Ref. [21] while the elastic form factors were taken from Ref. [22]. In Fig. 1 the distribution of events as a function of  $W^2$  is presented in comparison with the simulation. It is apparent that the shape of the simulated distribution agrees well with that of the data. It was found that the relative contaminations from the elastic and DIS regions in the yield of the resonance region range from 10% to 2% and from 7% to 16% respectively, as  $Q^2$  increases from  $1.2 \text{ GeV}^2$  up to  $12 \text{ GeV}^2$ .

Data were divided into six bins in  $Q^2$ , but only one bin in  $W^2$ . In each  $Q^2$ -bin the average longitudinal asymmetry  $A_1$  for virtual photoabsorption was calculated using the formula

$$A_1 = \frac{A_{\parallel}}{D} - \eta A_2, \quad (6)$$

where  $D$  and  $\eta$  are factors [13] that depend on kinematic variables. The quantity  $D$  depends also on  $R = \sigma_L/\sigma_T$ , the ratio of the absorption cross sections for longitudinal and transverse virtual photons.  $A_2$  is

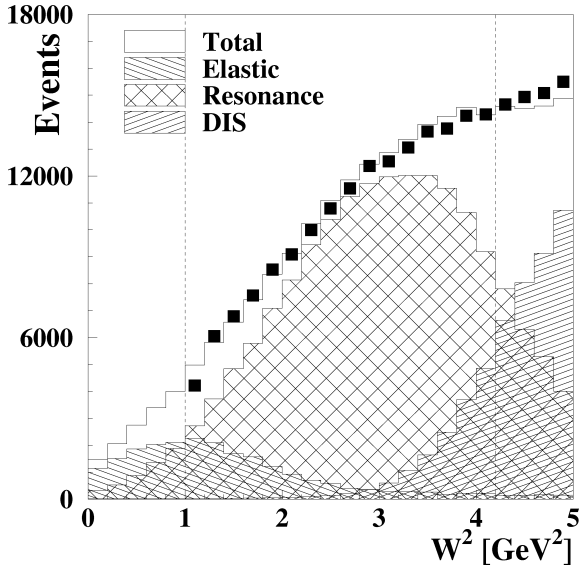


Fig. 1. Comparison of the event distribution for  $W^2 > 1 \text{ GeV}^2$  (squares) with the Monte-Carlo simulation (histogram). An overall normalisation factor was applied to the simulation to match the data. Also shown are the smeared distributions from the elastic, resonance and the deep inelastic regions. The vertical lines indicate the resonance region considered in the analysis.

related to longitudinal-transverse interference. The cross section asymmetry  $A_{\parallel}$  is given by:

$$A_{\parallel} = \frac{N^{\rightarrow} L^{\rightarrow} - N^{\rightarrow} L^{\leftarrow}}{N^{\rightarrow} L^{\rightarrow} + N^{\rightarrow} L^{\leftarrow}}. \quad (7)$$

Here,  $N^{\rightarrow}$  ( $N^{\leftarrow}$ ) is the number of scattered positrons for target spin parallel (anti parallel) to the beam spin orientation. The lifetime-corrected luminosities for

each target spin state are  $L^{\rightarrow(\leftarrow)}$  and  $L_P^{\rightarrow(\leftarrow)}$ , the latter being weighted by the product of beam and target polarisations. The cross-section asymmetry  $A_{\parallel}^{\text{res}}$  in the resonance region was corrected for contaminations originating from elastic and deep inelastic scattering, as discussed above. The asymmetry for the elastic contribution was taken from [22], while for the DIS region a parameterisation [23] based on world data has been used. Model dependent uncertainties due to these asymmetries and contributions from the Monte-Carlo simulation are negligible, as this correction is less than 5% for  $Q^2 < 5 \text{ GeV}^2$ . Radiative corrections were calculated using the codes described in Ref. [24], and were found not to exceed 2% of the asymmetry  $A_{\parallel}^{\text{res}}$ . The values for  $A_1^{\text{res}} + \eta A_2^{\text{res}}$  for the measured range of  $Q^2$  values are presented in Table 1. The asymmetry  $A_1$  was evaluated using Eq. (6) under the assumption that  $A_2 = 0.06$  [25] in the whole resonance region, and with an average depolarisation factor  $D$  weighted by the event distribution.

The contribution  $I_{\text{GDH}}^{\text{res}}$  of the resonance region to the GDH integral was determined in each  $Q^2$ -bin from the asymmetry  $A_1$ , according to Eq. (4), in which the integration limits were determined by the  $1 \text{ GeV}^2 < W^2 < 4.2 \text{ GeV}^2$  range. The unpolarised structure function  $F_1 = F_2(1 + \gamma^2)/(2x(1 + R))$  was calculated from a modification of the parameterisation of  $F_2$  given in Ref. [20] that accounts for nucleon-resonance excitation assuming  $R = \sigma_L/\sigma_T$  is constant and equal to 0.18 in the whole resonance region. It is worth noting that due to cancellation between the  $R$ -dependences of  $F_1$  and  $D$  at low  $y$ , the final result is insensitive to the choice of  $R$ . In the integration

Table 1

Results for  $A_1^{\text{res}} + \eta A_2^{\text{res}}$ , the resonance part ( $I_{\text{GDH}}^{\text{res}}$ ) to the GDH integral, and the total measured integral ( $I_{\text{GDH}}^{\text{meas}}$ ), as well as the full GDH integral ( $I_{\text{GDH}}$ ), including the unmeasured part. Errors represent the statistical uncertainty for  $A_1^{\text{res}} + \eta A_2^{\text{res}}$  and the statistical and the systematic uncertainties of the integrals

$\langle Q^2 \rangle [\text{GeV}^2]$	$A_1^{\text{res}} + \eta A_2^{\text{res}}$	$I_{\text{GDH}}^{\text{res}} [\mu\text{b}]$	$I_{\text{GDH}}^{\text{meas}} [\mu\text{b}]$	$I_{\text{GDH}} [\mu\text{b}]$
1.5	$0.71 \pm 0.16$	$21.4 \pm 5.2 \pm 4.1$	$37.9 \pm 5.5 \pm 5.1$	$41.2 \pm 5.5 \pm 5.1$
2.1	$0.77 \pm 0.18$	$10.3 \pm 2.5 \pm 2.0$	$24.3 \pm 2.8 \pm 2.9$	$27.8 \pm 2.8 \pm 2.9$
2.7	$0.74 \pm 0.22$	$4.9 \pm 1.5 \pm 0.9$	$17.5 \pm 1.8 \pm 1.8$	$21.0 \pm 1.8 \pm 1.8$
3.5	$0.79 \pm 0.22$	$2.4 \pm 0.7 \pm 0.4$	$13.0 \pm 1.0 \pm 1.2$	$16.5 \pm 1.0 \pm 1.2$
4.5	$0.97 \pm 0.29$	$1.3 \pm 0.4 \pm 0.2$	$8.8 \pm 0.6 \pm 0.8$	$12.3 \pm 0.6 \pm 0.8$
6.6	$0.55 \pm 0.23$	$0.08 \pm 0.03 \pm 0.01$	$5.3 \pm 0.3 \pm 0.5$	$8.6 \pm 0.3 \pm 0.5$



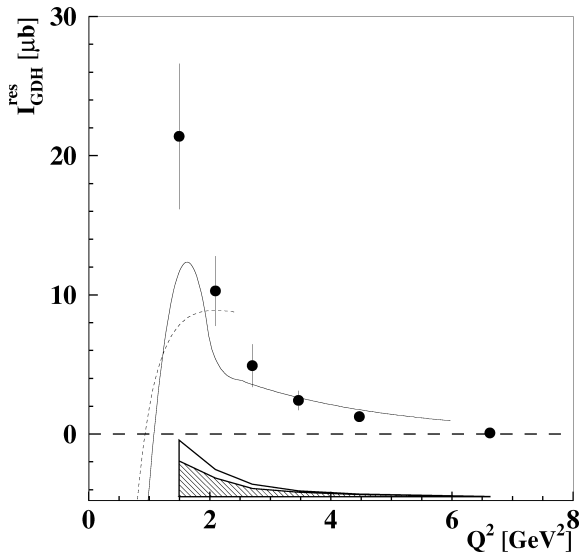


Fig. 2. The GDH integral as a function of  $Q^2$  for the region  $1.0 \text{ GeV}^2 < W^2 < 4.2 \text{ GeV}^2$ . The error bars show the statistical uncertainties. The white and the hatched bands represent the systematic uncertainties with and without the  $A_2$  uncertainty contribution. The dashed [9] and the solid [10] curves are predictions based on a  $Q^2$ -evolution of nucleon-resonance amplitudes.

the  $W^2$  dependence of the integrand  $F_1/K$  within the individual bins was fully accounted for.

The results for  $I_{\text{GDH}}^{\text{res}}$  are presented in Fig. 2. The integral strongly decreases with  $Q^2$  over the entire measured range. The magnitude of the systematic uncertainty is indicated by the band. The dominant contribution to the systematic uncertainty is due to the lack of knowledge of  $A_2$ . This contribution (up to 15%) was evaluated from the total error quoted for a measurement in the resonance region:  $A_2 = 0.06 \pm 0.16$  [25]. This range is consistent with two other possible assumptions for  $A_2$ :  $A_2 = 0$ , or  $A_2 = 0.53 M_x / \sqrt{Q^2}$ , which describes the behaviour in the deep inelastic region [18]. Other contributions are uncertainties from the beam and target polarisations (5.3%), from the spectrometer geometry (2.5%), from the combined smearing and radiative effects (up to 10%) and from the knowledge of  $F_2$  (2%). The smearing contribution to the systematic uncertainty was evaluated by comparing simulated results from two very different assumptions for  $A_1$ : a power law ( $A_1 = x^{0.727}$ ) that smoothly extends the DIS behaviour for the asymmetry into the resonance region [23],

and a step function ( $A_1 = -0.5$  for  $W^2 < 1.8 \text{ GeV}^2$  and  $A_1 = +1.0$  for  $1.8 \text{ GeV}^2 < W^2 < 4.2 \text{ GeV}^2$ ) that is suggested by the hypothesis of the possible dominance of the  $P_{33}$ -resonance at low  $W^2$  and of the  $S_{11}$ -resonance at higher  $W^2$ .

The results for  $I_{\text{GDH}}^{\text{res}}$  are compared in Fig. 2 with two predictions for the contribution of nucleon-resonance excitation to the integral defined in Eq. (2). Burkert and Li [9] parameterised the experimental  $Q^2$ -evolution of the main nucleon resonances ( $P_{33}(1232)$ ,  $P_{11}(1440)$ ,  $S_{11}(1535)$ ,  $D_{13}(1520)$ ,  $F_{15}(1680)$ ), and assumed single-quark transitions to evaluate the contributions from other resonances. Aznauryan [10] described the resonance excitation in the approximation of infinitely narrow resonances, and included a contribution from one-pion exchange in the near-threshold region. Both models predict a sudden decline in  $I_{\text{GDH}}^{\text{res}}$  as  $Q^2$  falls below  $1.5 \text{ GeV}^2$ , due to a large negative contribution at low  $Q^2$  by the helicity structure of the  $P_{33}$ -resonance. At higher  $Q^2$  the  $P_{33}$ -resonance magnetic form factor strongly decreases with increasing  $Q^2$ , and the positive contribution to  $I_{\text{GDH}}^{\text{res}}$  arising from the excitation of higher-mass resonances becomes dominant. Neither of these models includes the nonresonant multihadron production channels, which should provide an additional positive contribution for the region  $W^2 \leq 4.2 \text{ GeV}^2$ . Comparison with the data suggests that for  $Q^2 \simeq 1.5 \text{ GeV}^2$ , the resonance excitation models are not sufficient to fully explain the experimental result for  $I_{\text{GDH}}^{\text{res}}$ . Other predictions exist for the resonance-excitation contribution to generalised GDH integrals, but they are limited to regions of lower  $Q^2$  [26].

To complete the evaluation of the full integral  $I_{\text{GDH}}$ , data from the DIS region ( $4.2 \text{ GeV}^2 < W^2 < 45 \text{ GeV}^2$ ) were reanalysed in the same  $Q^2$ -bins as for the kinematically more restricted resonance region, following the procedure described in a previous HERMES publication [13]. A total of 1.52 million events were selected in this  $W^2$ -range. The systematic uncertainty for this region is the same as published in [13,18]. The systematic uncertainty on  $A_2$  in DIS region does not contribute significantly.

In Table 1 the resonance-region contribution  $I_{\text{GDH}}^{\text{res}}$ , the integrals  $I_{\text{GDH}}^{\text{meas}}$  in the full measured region and the full GDH integrals are reported. The latter was calculated in each  $Q^2$ -interval by adding to  $I_{\text{GDH}}^{\text{meas}}$  an

estimate of the unmeasured contribution for  $W^2 > 45 \text{ GeV}^2$ . This was calculated using a multiple-reggeon exchange parameterisation [3] for  $\sigma_{1/2}(\nu, Q^2) - \sigma_{3/2}(\nu, Q^2)$  at high energy, and amounted to about  $3.5 \mu\text{b}$  for all  $Q^2$ -bins. A parameterisation for  $g_1$  [27] based on a NLO-QCD analysis provided within 5% the same results as the multiple-reggeon exchange analysis. This difference was taken as the systematic uncertainty of the high-energy contribution. It is worth noting that for the low  $Q^2$ -bins, both the statistical and the systematic uncertainties of  $I_{\text{GDH}}$  are dominated by those from the resonance region. In this region, these uncertainties are large due to the smallness of  $D$  and to the large size of  $\eta$ , respectively.

In Fig. 3a the total GDH integral is shown together with the partial integrals for  $W^2 < 4.2 \text{ GeV}^2$  and for  $W^2 < 45 \text{ GeV}^2$ . The contribution of the resonance region to the full GDH integral is small for  $Q^2$  values above  $3 \text{ GeV}^2$ .

Fig. 3a also shows a prediction [8] based on a  $Q^2$ -evolution of the two polarised structure functions  $g_1$  and  $g_2$ , without consideration of any explicit nucleon-resonance contribution. This prediction is in good agreement with the experimental data.

In the whole energy range,  $I_{\text{GDH}}$  is consistent within the uncertainties ( $\chi^2/N_{\text{df}} = 0.4$ ) with a simple  $1/Q^2$  power law. This is demonstrated in Fig. 3b where the results for  $I_{\text{GDH}}$  are multiplied by  $Q^2/(16\pi^2\alpha)$ . In the deep inelastic limit, this quantity is equivalent to  $\Gamma_1$  (see Eq. (5)). The present results are in agreement with the measurements of  $\Gamma_1 = 0.120 \pm 0.016$  at  $Q^2 = 10 \text{ GeV}^2$  [28] and  $\Gamma_1 = 0.129 \pm 0.010$  at  $Q^2 = 5 \text{ GeV}^2$  [25]. In addition, values of  $\Gamma_1$  extracted from the present data are also consistent with a measurement of  $\Gamma_1 = 0.104 \pm 0.017$  at  $Q^2 = 1.2 \text{ GeV}^2$  [25] in which the structure function  $g_1$  was measured in the resonance region.

In summary, the  $Q^2$ -dependence of the generalised Gerasimov–Drell–Hearn integral for the proton was determined for the first time in both the resonance and the deep inelastic regions, covering the  $Q^2$ -range from  $1.2$  to  $12 \text{ GeV}^2$ . In the resonance region, the data suggest that for  $Q^2 \geq 1.5 \text{ GeV}^2$ , existing resonance-excitation models are not sufficient to fully explain the experimental result for  $I_{\text{GDH}}^{\text{res}}$ . Above  $Q^2 = 3 \text{ GeV}^2$  the DIS contribution to the generalised GDH integral is dominant. The  $Q^2$ -behaviour of  $I_{\text{GDH}}$  suggests that there are no large effects from either resonances or

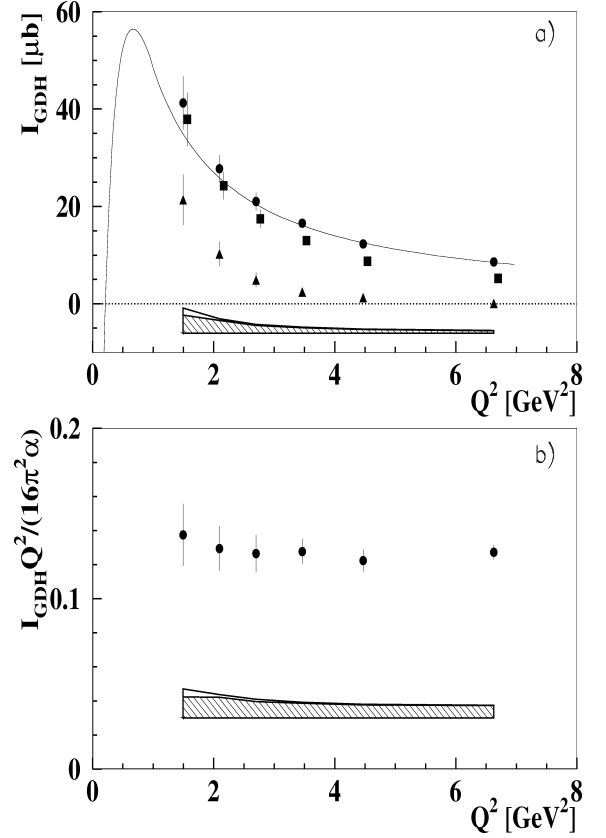


Fig. 3. (a)  $I_{\text{GDH}}$  as a function of  $Q^2$  for various upper limits of integration:  $W^2 \leq 4.2 \text{ GeV}^2$  (triangles),  $W^2 \leq 45 \text{ GeV}^2$  (squares), and the total integral  $I_{\text{GDH}}$  (circles). The squares have been slightly shifted to make them more visible. The curve is the Soffer–Teryaev model [8] for the total integral. (b)  $I_{\text{GDH}} Q^2 / (16\pi^2\alpha)$  as a function of  $Q^2$ . For both panels, the error bars show the statistical uncertainties, and the white and the hatched bands at the bottom represent the systematic uncertainties for the total integral with and without the  $A_2$  contribution.

nonleading twist, and indicates that the sign change of  $I_{\text{GDH}}$  to meet the real photon limit occurs at  $Q^2$  lower than  $1.2 \text{ GeV}^2$ .

### Acknowledgements

We thank S. Gerasimov, V. Burkert and I.G. Aznauryan for useful discussions and I.G.A. for providing the curves of her calculations. We gratefully acknowledge the DESY management for its support, the staffs at DESY and the collaborating institutions for their



significant effort, and our funding agencies for financial support.

## References

- [1] S.B. Gerasimov, *Sov. J. Nucl. Phys.* 2 (1966) 430;  
S.D. Drell, A.C. Hearn, *Phys. Rev. Lett.* 16 (1966) 908.
- [2] I. Karliner, *Phys. Rev. D* 7 (1973) 2717;  
R.L. Workman, R.A. Arndt, *Phys. Rev. D* 45 (1992) 1789;  
A.M. Sandorfi, C.S. Whisnant, M. Khandaker, *Phys. Rev. D* 50 (1994) R6681;  
D. Drechsel, G. Krein, *Phys. Rev. D* 58 (1998) 116009.
- [3] N. Bianchi, E. Thomas, *Phys. Lett. B* 450 (1999) 439.
- [4] A. Thomas, *Nucl. Phys. B* 79 (1999) 591c.
- [5] R. Pantförder, PhD Thesis, Universität Bonn, 1998, BONN-IR-98-06, hep-ph/9805434, and references therein.
- [6] F.J. Gilman, *Phys. Rev.* 167 (1968) 1365.
- [7] M. Anselmino, B.L. Ioffe, E. Leader, *Yad. Fiz.* 49 (1989) 214.
- [8] J. Soffer, O.V. Teryaev, *Phys. Rev. D* 51 (1995) 25;  
J. Soffer, O.V. Teryaev, *Phys. Rev. Lett.* 70 (1993) 3373.
- [9] V. Burkert, Z.J. Li, *Phys. Rev. D* 47 (1993) 46;  
V. Burkert, B. Ioffe, *JETP* 105 (1994) 1153.
- [10] I.G. Aznauryan, *Phys. At. Nucl.* 58 (1995) 1014, and private communication.
- [11] J. Edelmann, G. Piller, N. Kaiser, W. Weise, Technische Universität München T39-99-19, hep-ph/9909524.
- [12] V. Bernard, N. Kaiser, U. Meissner, *Phys. Rev. D* 48 (1993) 3062;  
X. Ji, C.-W. Kao, J. Osborne, *Phys. Lett. B* 472 (2000) 1.
- [13] HERMES Collaboration, K. Ackerstaff et al., *Phys. Lett. B* 444 (1998) 531.
- [14] D.P. Barber et al., *Nucl. Instrum. Methods A* 329 (1993) 79;  
A. Most, in: C.W. de Jager et al. (Eds.), *Proc. 12th Int. Symp. on High-Energy Spin Physics*, Amsterdam, The Netherlands, World Scientific, 1997, p. 800.
- [15] J. Stewart, in: R.J. Holt, M.A. Miller (Eds.), *Proc. of the Workshop Polarised Gas Targets and Polarised Beams*, Urbana-Champaign, USA, AIP Conf. Proc., Vol. 421, 1997, p. 69.
- [16] F. Stock et al., *Nucl. Instrum. Methods A* 343 (1994) 334.
- [17] B. Braun, in: R.J. Holt, M.A. Miller (Eds.), *Proc. of the Workshop Polarised Gas Targets and Polarised Beams*, Urbana-Champaign, USA, AIP Conf. Proc., Vol. 421, 1997, p. 156.
- [18] HERMES Collaboration, A. Airapetian et al., *Phys. Lett. B* 442 (1998) 484.
- [19] HERMES Collaboration, K. Ackerstaff et al., *Nucl. Instrum. Methods A* 417 (1998) 230.
- [20] A. Bodek et al., *Phys. Rev. D* 20 (1979) 1471.
- [21] NMC Collaboration, P. Amaudruz et al., *Nucl. Phys. B* 371 (1992) 3.
- [22] S.I. Bilen'kaya, L.I. Lapidus, S.M. Bilen'kii, Yu.M. Kazari-nov, *Zh. Eksp. Teor. Fiz. Pis'ma* 19 (1974) 613.
- [23] A.P. Nagaitsev et al., JINR Rapid Communication, July 1995, N3(71)-95,59.
- [24] I.V. Akushevich, N.M. Shumeiko, *J. Phys. G* 20 (1994) 513;  
I. Akushevich et al., *Comput. Phys. Commun.* 104 (1997) 201.
- [25] E143 Collaboration, K. Abe et al., *Phys. Rev. D* 58 (1998) 112003.
- [26] W.X. Ma, D.H. Lu, A.W. Thomas, Z.P. Li, *Nucl. Phys. A* 635 (1998) 497;  
D. Drechsel, S.S. Kamalov, G. Krein, L. Tiator, *Phys. Rev. D* 59 (1999) 094021;  
Y. Dong, *Phys. Lett. B* 425 (1998) 177;  
O. Scholten, A.Y. Korshin, *Eur. Phys. J. A* 6 (1999) 211.
- [27] M. Glück et al., *Phys. Rev. D* 53 (1996) 4775.
- [28] SMC Collaboration, B. Adeva et al., *Phys. Rev. D* 58 (1998) 112001.